

- Intrapixel, interpixel and jitter
- Evolution of detector within an observing visit
- Variation of the instrument response with time

Minimizing :

OBS(t)-TRANSIT_MODEL(planet) \otimes INSTRUMENT_MOD(t)

- Basis of orthogonal functions
- Independant component analysis
- Other

EChO requirement

Stability 10⁻⁴ over 10 hours

FINESSE study:

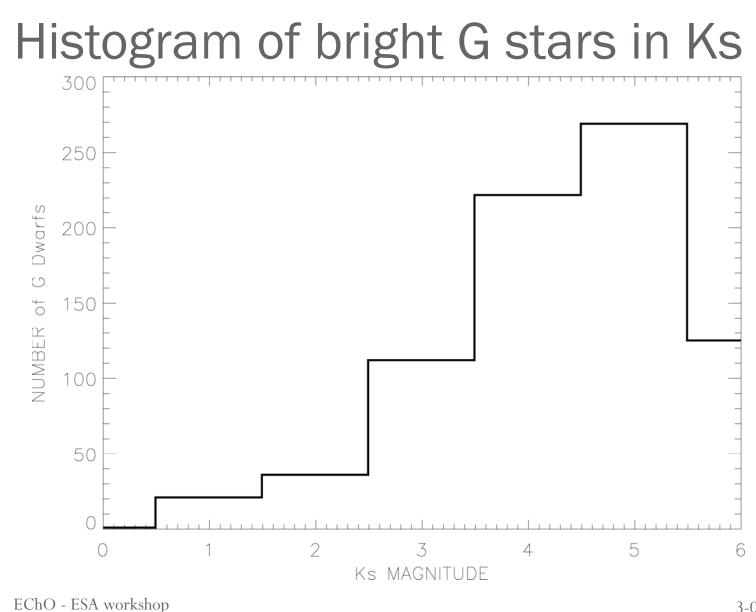
- 1 calibrator after any transit observation (1h)
- Every week, 1 calibrator for 8 hours

EChO (a proposition):

- 1 calibrator (1h) every \sim 36 hours
- 1 calibrator (10h) every ~ 10 days

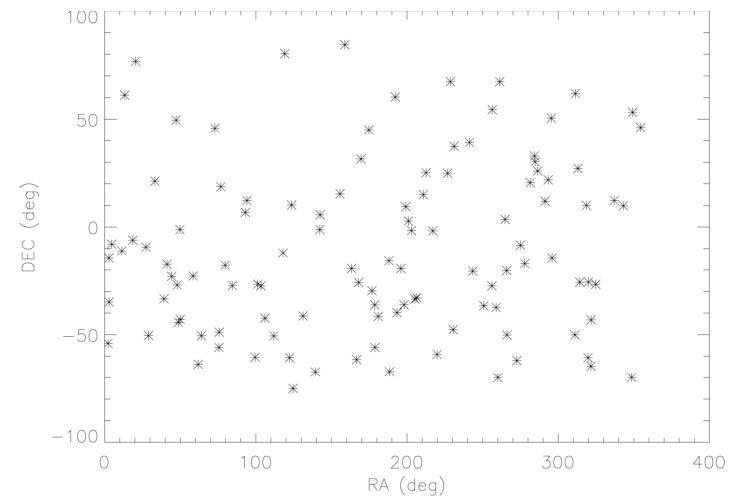
Getting standard candles stable at few 10⁻⁵

- G stars are the best candles you can dream of.
- Using Kepler, Ciardi et al. 2012
 - G dwarf stars are the more stable
 - Variability floor 10⁻⁵ over 30 days for 70 % of G dwarfs



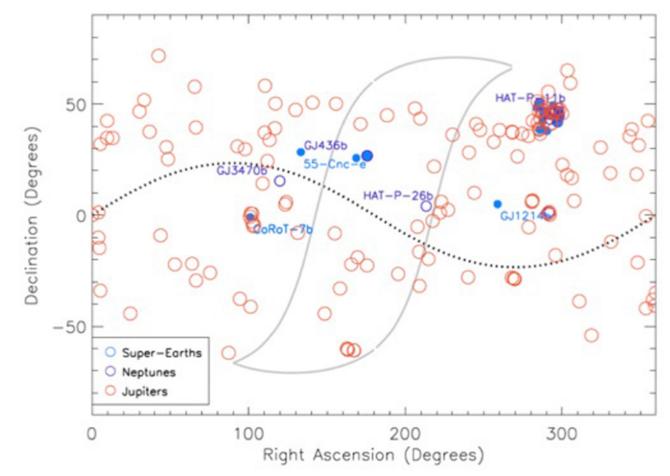
3-07-2013

Distribution of bright G stars on the sky



3-07-2013

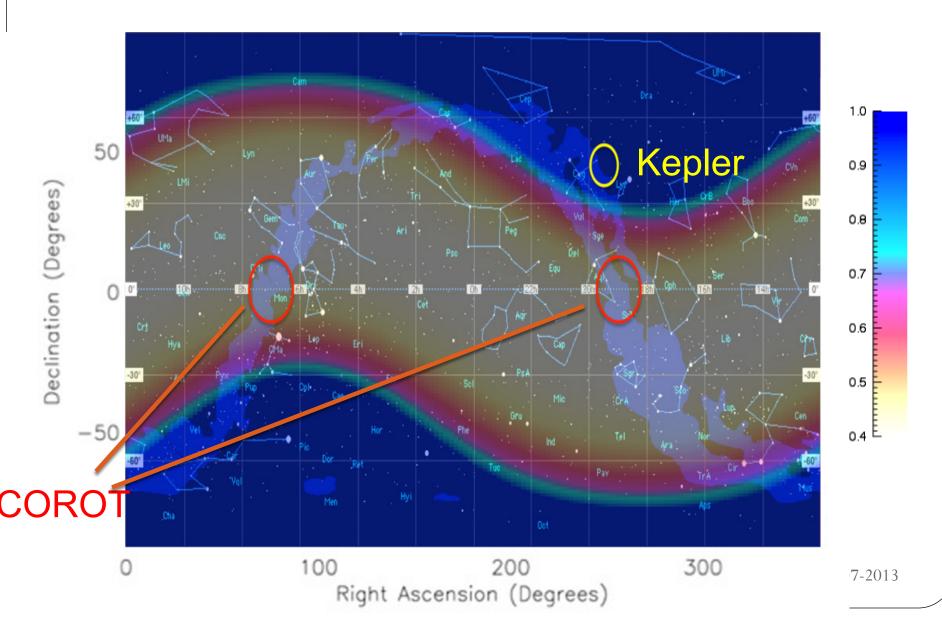
Visibility of known transiting extrasolar planets



3-07-2013

6

EChO visibility



Remarks from scheduling exercice

Constraint from scheduling exercise :

- 1 hour observation every \sim 30 hours is fairly easy to schedule
- 10 hours every 10 days is a constraint on schedule
- 8 hours is easier to schedule than 10 hours
- On year 1, calibrations are more important but harder to schedule.

Concluding remarks

Constraining your instrument model:

• Using few bright stable G dwarf stars in continuous viewing zone

- Stability information from TESS
- Dedicated survey by Cheops to find best calibrators
- Spectropolarimetry + Ca HK lines survey to monitor activity

• How long and how often ?

- Conservative on year 1 to learn the instrument.
- (1 hour calib every ~36 hours, 10 hours calib every ~10 days)