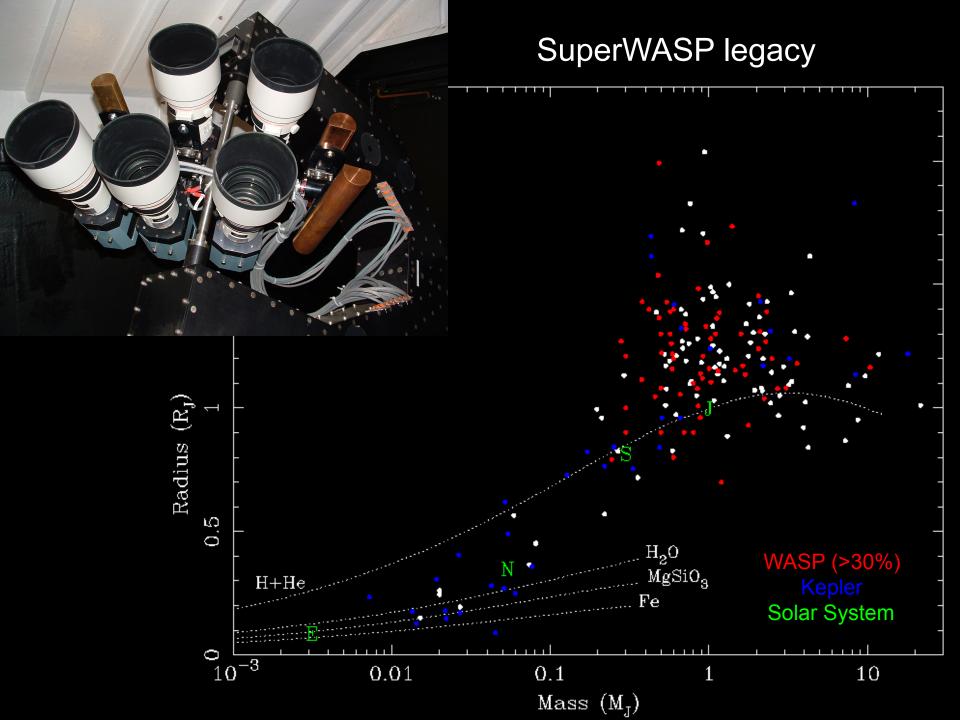
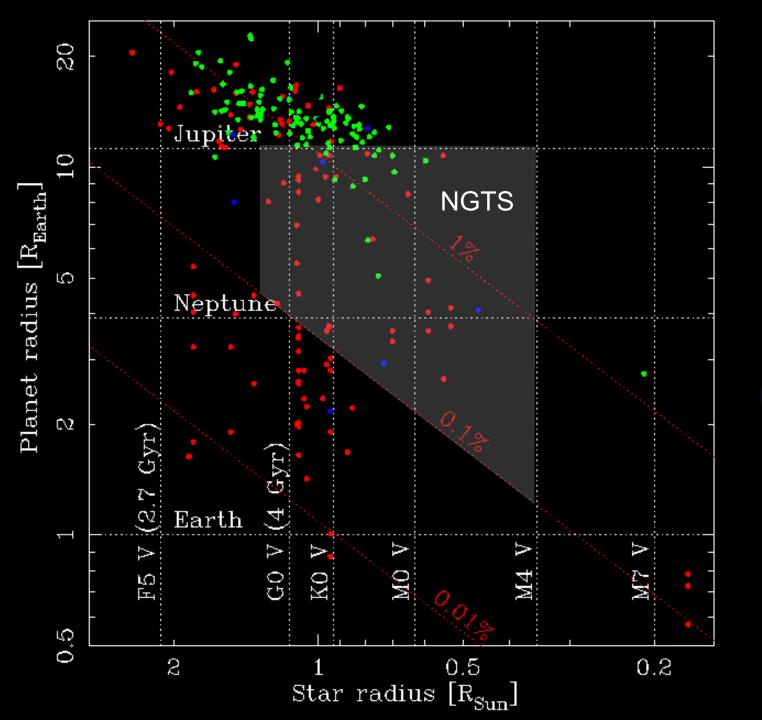
### NGTS: Next-Generation Transit Survey

Peter Wheatley, University of Warwick

Belfast, Cambridge, DLR Berlin, Geneva, Leicester, Warwick



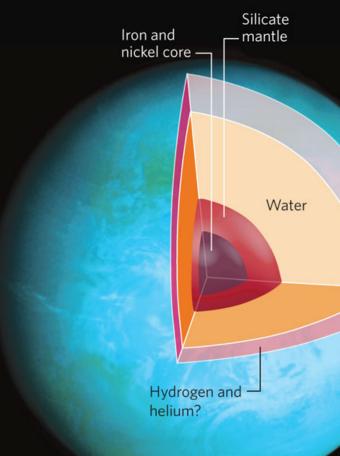




#### Transit discovery space

Ground-based Space-based Radial velocity Key science goals:

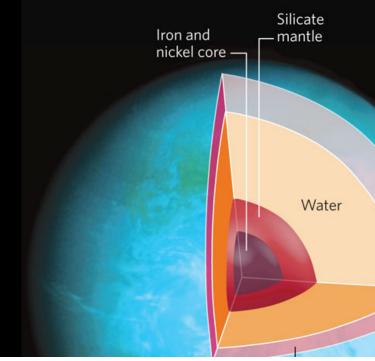
- Statistical sample (>100) of Neptunes and Super-Earths with measured mass, density, orbital separation
- Very bright systems for atmosphere studies with VLT, E-ELT, JWST, EChO

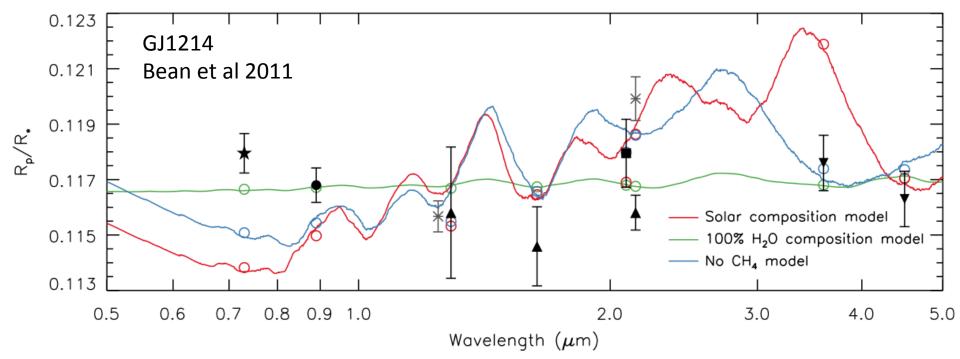




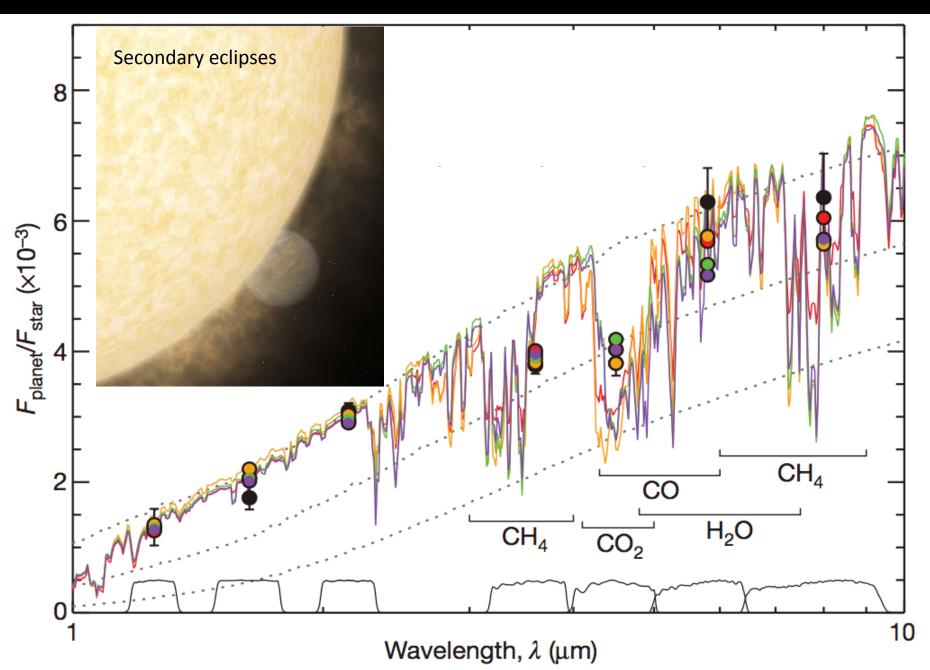
Key science goals:

- Statistical sample (>100) of Neptunes and Super-Earths with measured mass, density, orbital separation
- Very bright systems for atmosphere studies with VLT, E-ELT, JWST, EChO



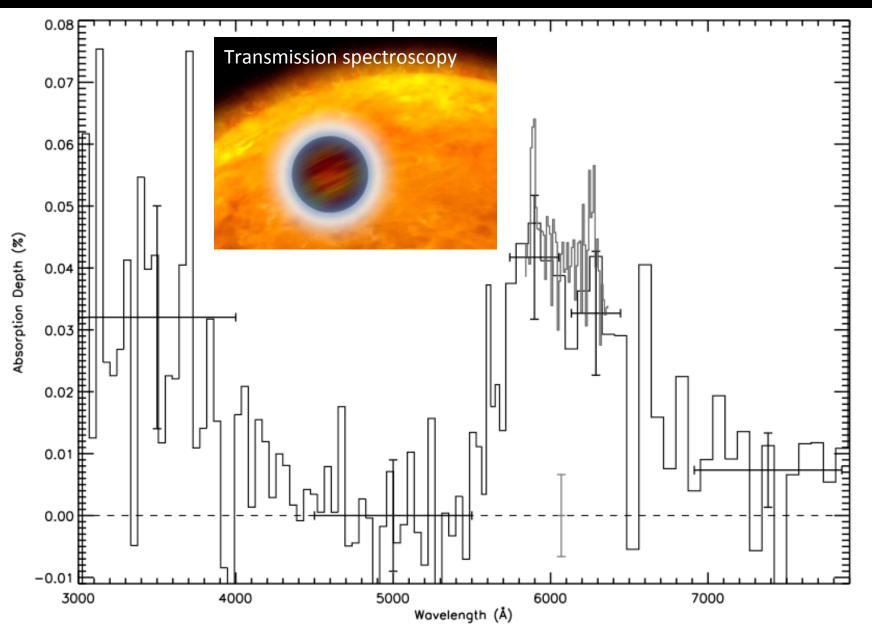


#### Carbon rich atmosphere in WASP-12b



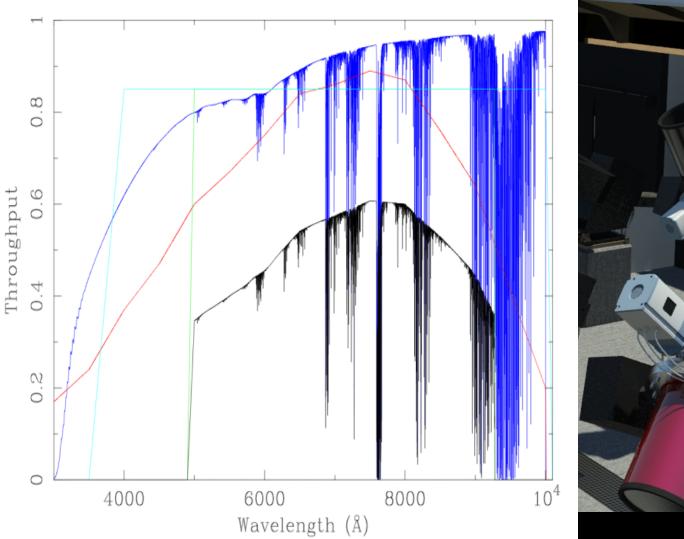
#### Rayleigh scattering and sodium in a planetary atmosphere

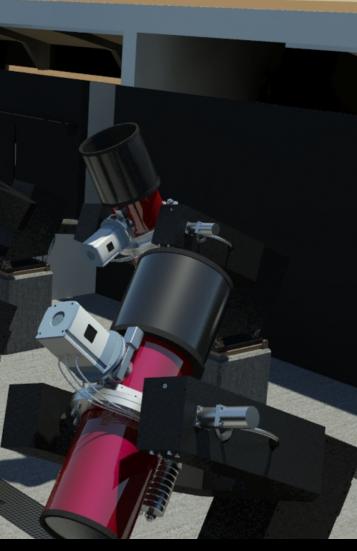
Sing et al. 2008



# NGTS Design

12 x 20cm f/2.8 telescopes 110 sq deg 600-900nm

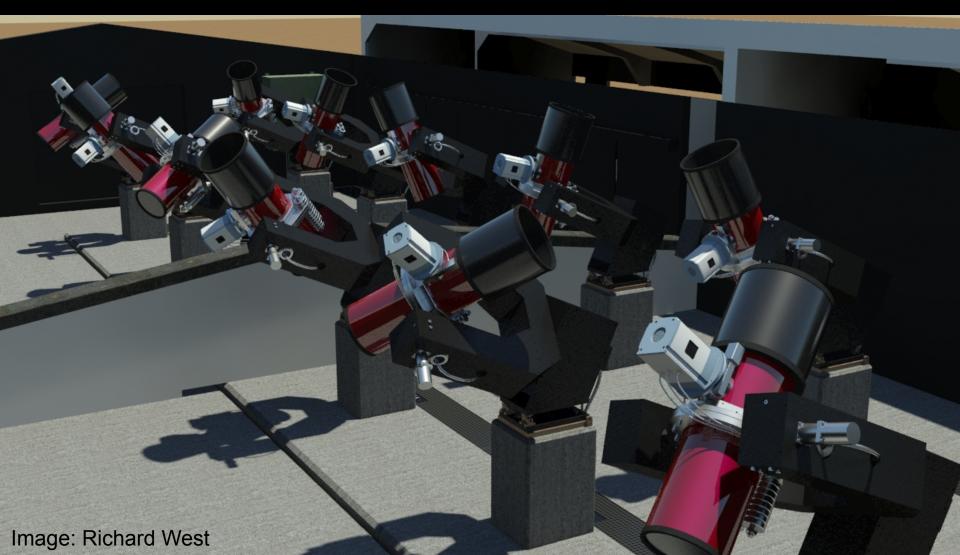




#### **NGTS Facility**

#### 12 x 20cm f/2.8 telescopes 96 sq deg field of view deep depletion CCDs (550-900nm)

£2m capital investment by Warwick, Leicester, QUB, Geneva, DLR Berlin, Cambridge



#### NGTS Site: ESO Paranal observatory, Chile

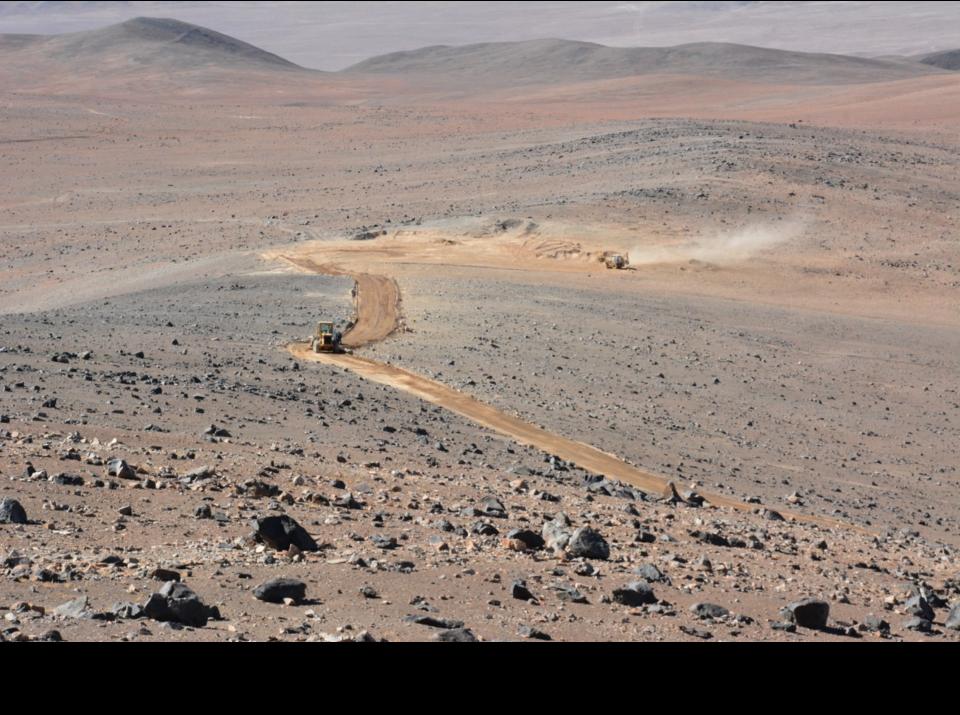
VLT

Construction 2013 Operations from 2014

ESO providing public data archive

VISTA

NGTS 🥢

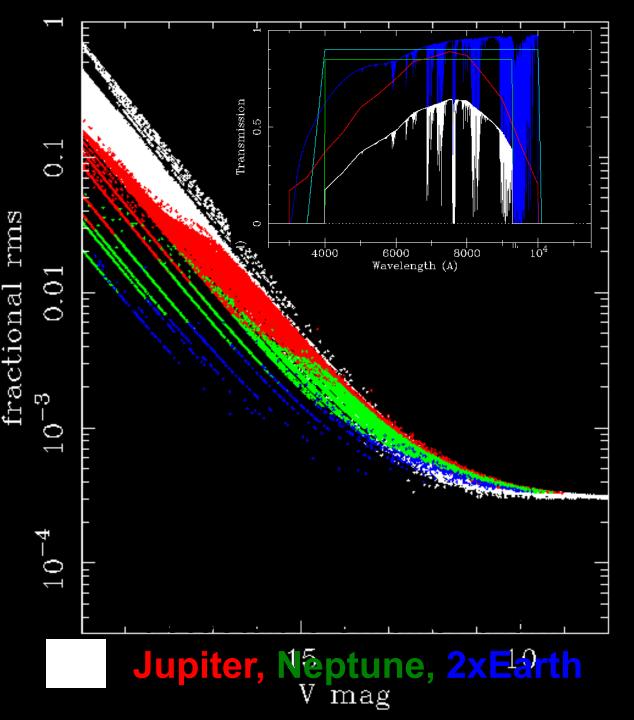


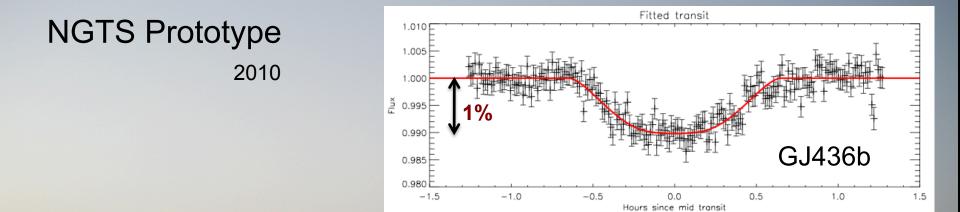


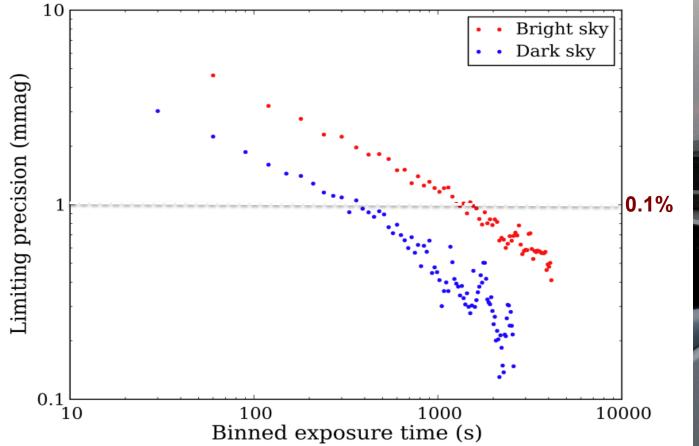


## NGTS simulations

- >100 RV-confirmed Neptunes and super-Earths giving bulk composition
- ∼20 Neptunes and super-Earths for detailed atmospheric characterisation (I<11.5)</li>







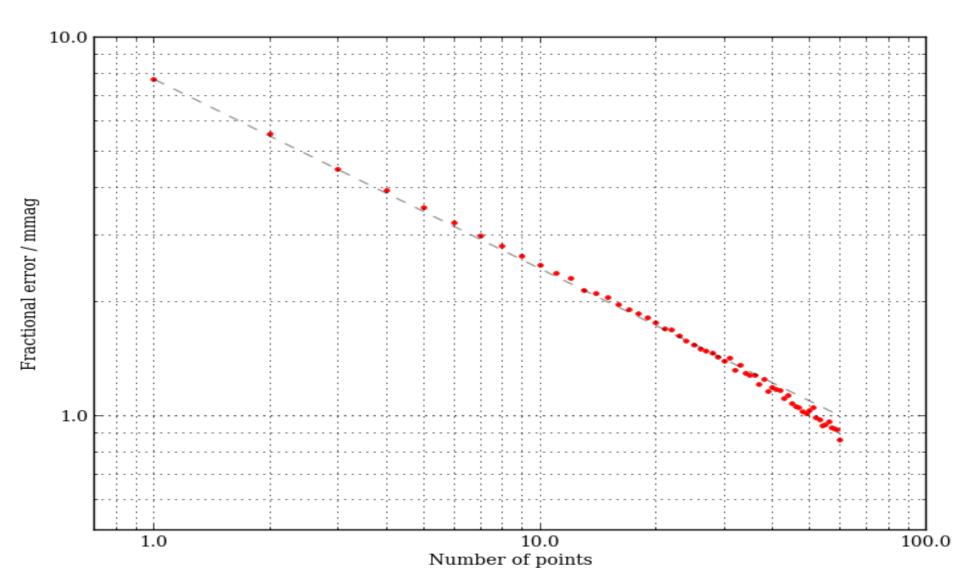


System testing in Geneva August 2012

### System testing in Geneva

August 2012

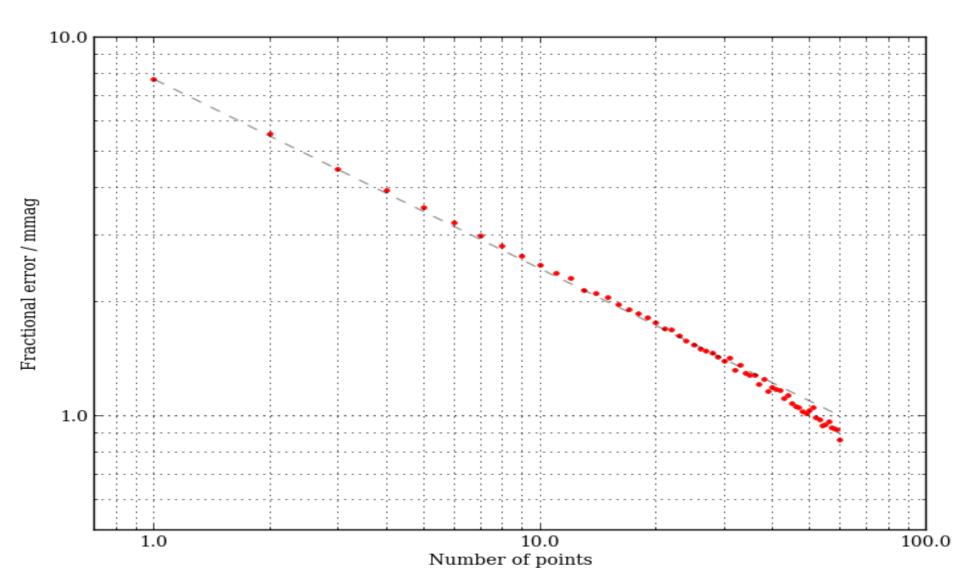




### System testing in Geneva

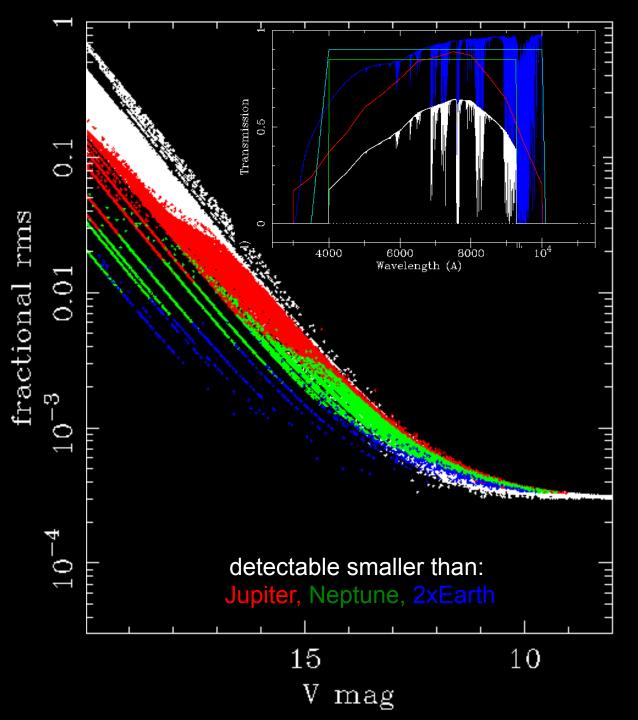
August 2012





# NGTS planet catch simulations

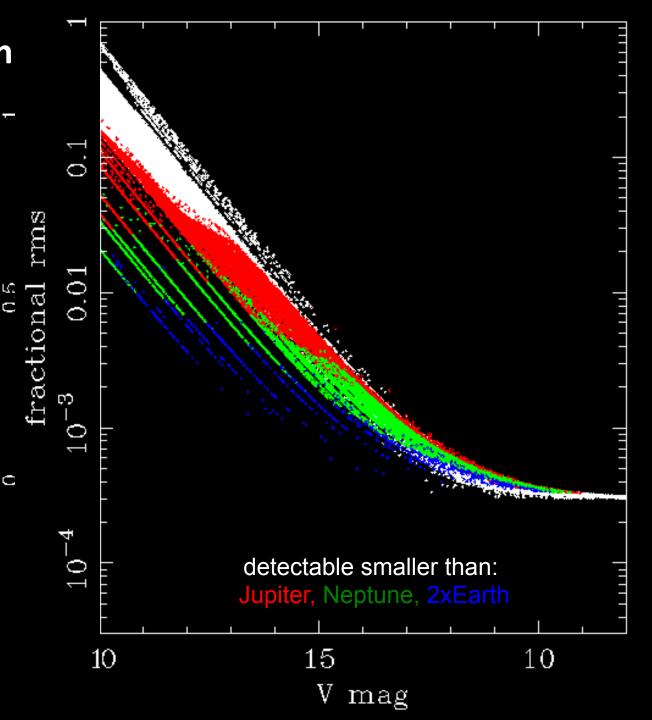
- Instrumental throughput
- Sky transmission
- Real sky background
- Galactic model
- Real stellar spectra
- Scintillation
- Red noise floor
- Planet distributions and occurrence rates from Kepler
- Weather at Paranal
- Period window function
- Transit detection
  probability
- Radial velocity sensitivity



# NGTS planet catch simulations

lransmission

- Instrumental throughput
- Sky transmission
- Real sky background
- Galactic model
- Real stellar spectra
- Scintillation
- Red noise floor
- Planet distributions and occurrence rates from Kepler
- Weather at Paranal
- Period window function
- Transit detection probability
- Radial velocity sensitivity



# NGTS planet catch simulations

90G

200

Ð

800

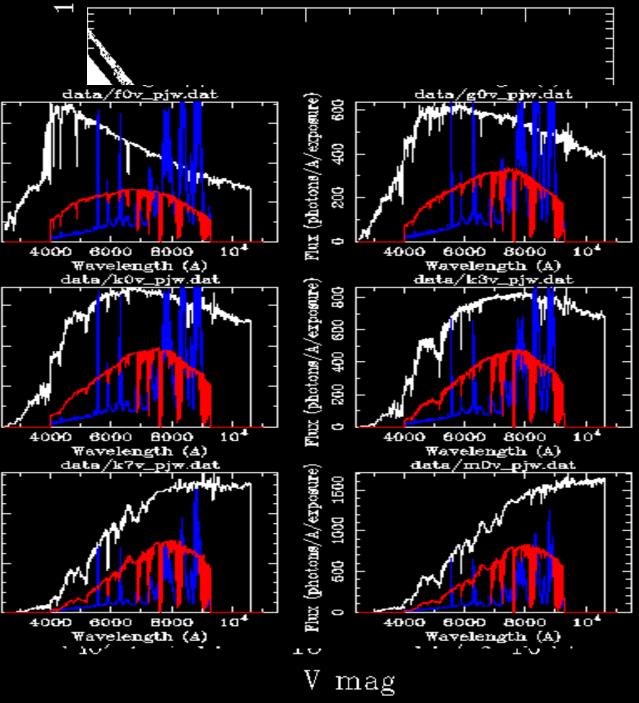
200

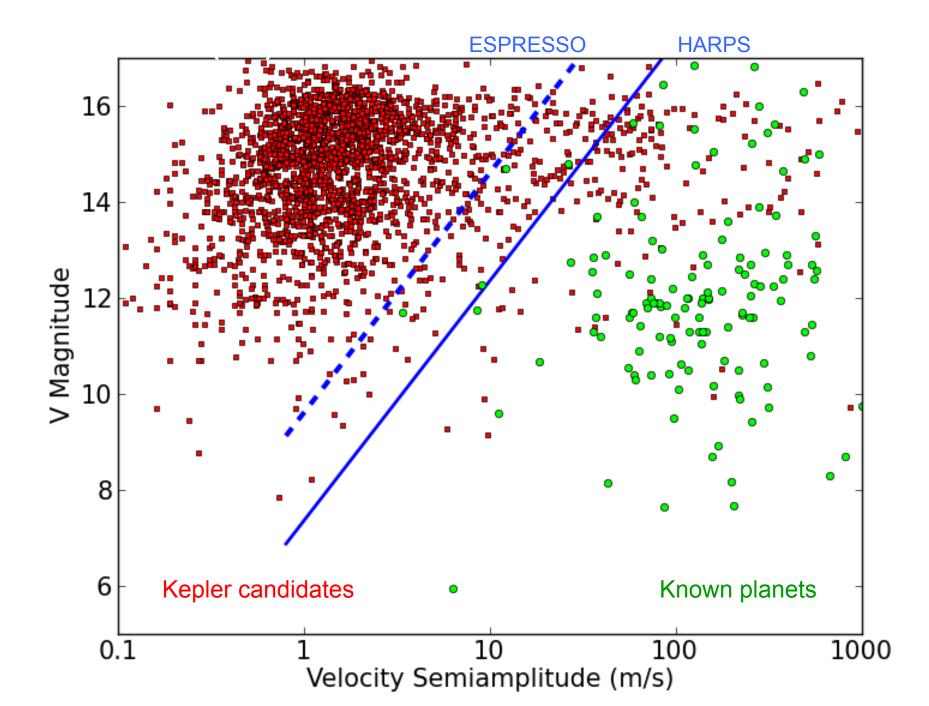
000

0

smissi

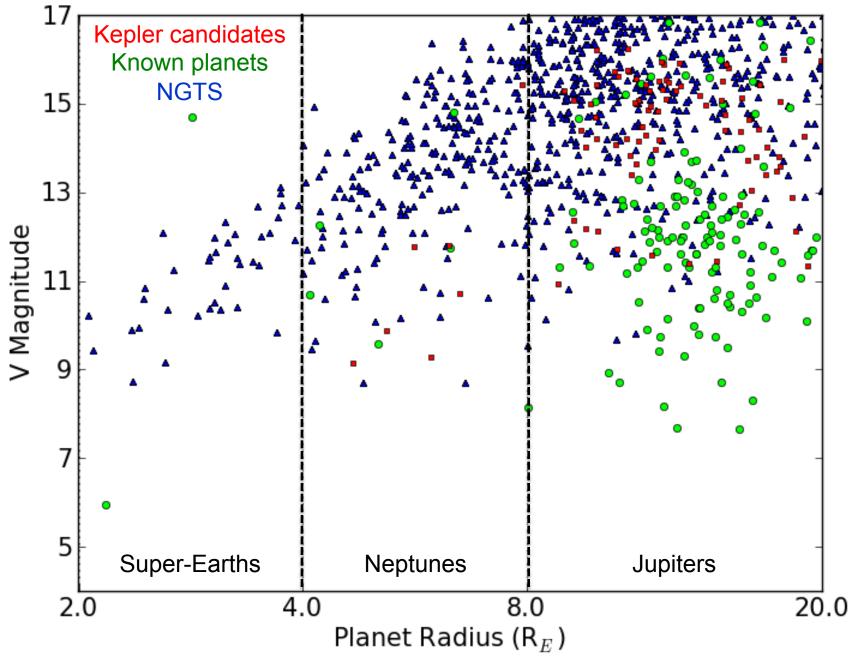
- Instrumental throughput ន្ន
- Sky transmission
- Real sky background
- Galactic model
- Real stellar spectra
- Scintillation
- Red noise floor
- Planet distributions and occurrence rates from Kepler
- Weather at Paranal
- Period window function
- Transit detection probability
- Radial velocity sensitivity

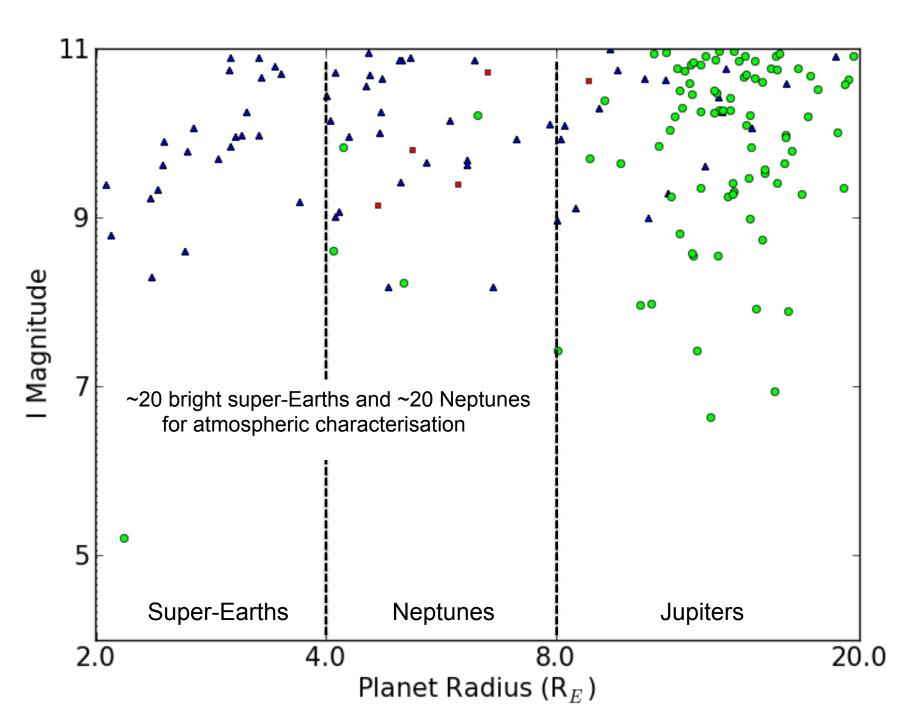




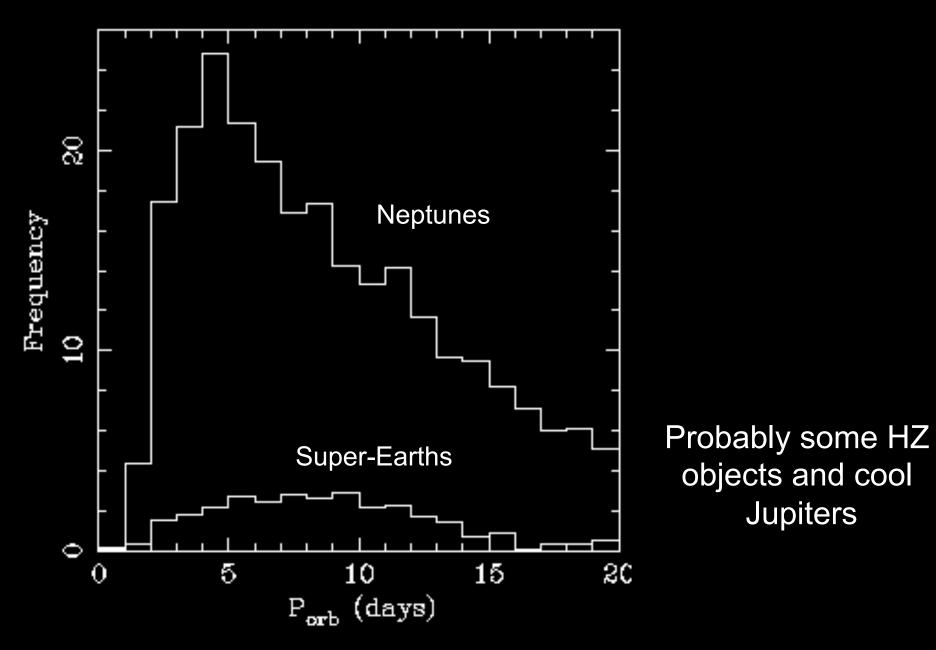
~40 RV-confirmed super-Earths and

~230 Neptunes for bulk composition





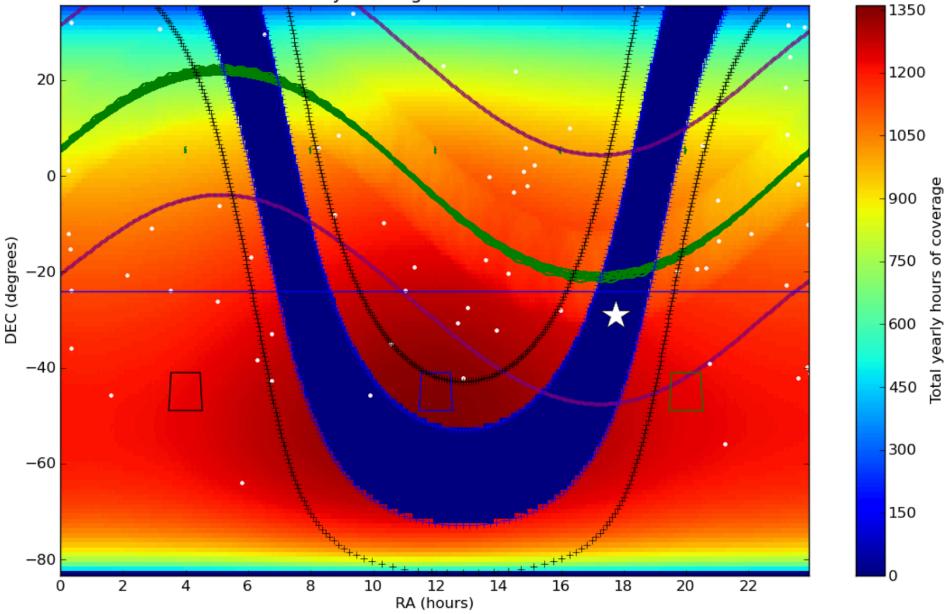
### **Period distributions**

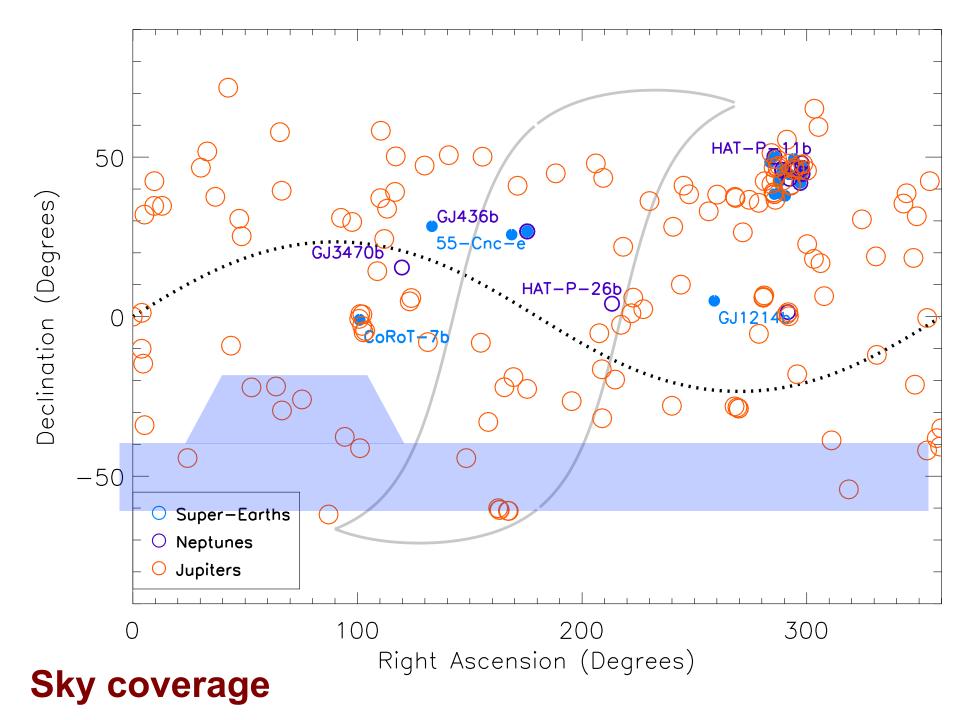


## Simulated planet catch (Q3)

- 10 hours / object with HARPS / HARPS-N
  - 37 NGTS candidate Neptunes can be confirmed, cf 7 from Kepler
  - Together raising total known to 48
  - Neither mission sensitive to super-Earths (1 predicted from NGTS)
- 20 hours / object with HARPS-N
  - Kepler Neptunes increased to only 21, and only 1 super-Earth
- 10 hours / object with ESPRESSO (from 2016)
  - NGTS can confirm 231 Neptunes and 39 super-Earths
  - NGTS bright sample consists of 25 Neptunes and 23 super-Earths

#### NGTS sky coverage for 2012 at Paranal





#### Wheatley et al., 2013, arXiv:1302.6592 www.ngtransits.org

Belfast Berlin Cambridge Geneva Leicester Warwick

